



# Trends in High Performance Computing, and their Impact on Astrophysical Data Processing

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# $C^3$ at LBNL

- Focused on computational challenges (simulation and data processing) relevant to cosmology (CMB, SN, BAO, ...)
- Tight connection to DOE computing facilities: Cray XT5 (40K cores), Cray XE6 (150K cores), Cloud computing platform, GPU test cluster, science gateways, etc.
- For >10 years, we have coordinated CPU allocations for CMB telescopes (funded by NASA, NSF, etc).
- Involved in building software infrastructure for future experiments and future architectures: algorithm scaling, data management, etc.

# *High Performance Computing*

## **For the purposes of this talk, everything that needs a machine room:**

- Traditional Clusters (PCs interconnected with ethernet, infiniband, etc)
- Supercomputers (lightweight nodes with infiniband or custom interconnect)
- Cloud computing platforms (EC2, Eucalyptus)
- Large shared memory machines (NUMA architectures)

# *HPC in 10 Years*

## **Hard to predict, but driven by trends:**

- Still using silicon, and still tracking Moore's law for transistor counts.
- Computing centers have limited electrical capacity for power and cooling.
- Packing transistors into traditional CPU cores requires even more transistors for “overhead”-diminishing returns.
- Market forces (follow the money)

# Moore's Law

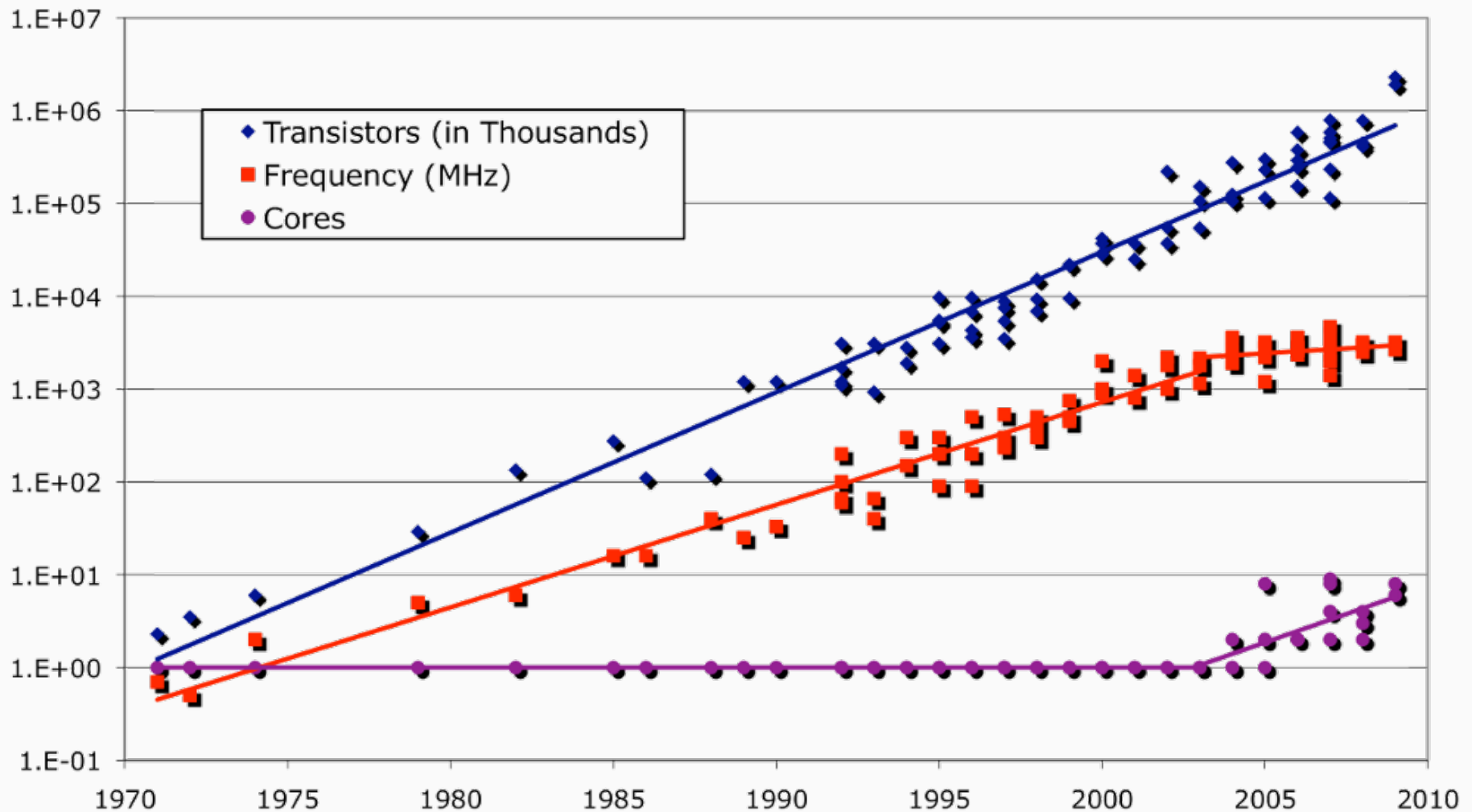


Figure by Kathy Yelick, data from Kunle Olukotun, Lance Hammond, Herb Sutter, Burton Smith, Chris Batten, and Krste Asanovic

# *Rise of Many-core Systems*

## **Focus is on Flops per Watt:**

- Clock rates constant or decreasing.

# Clock Rates and Power Scaling

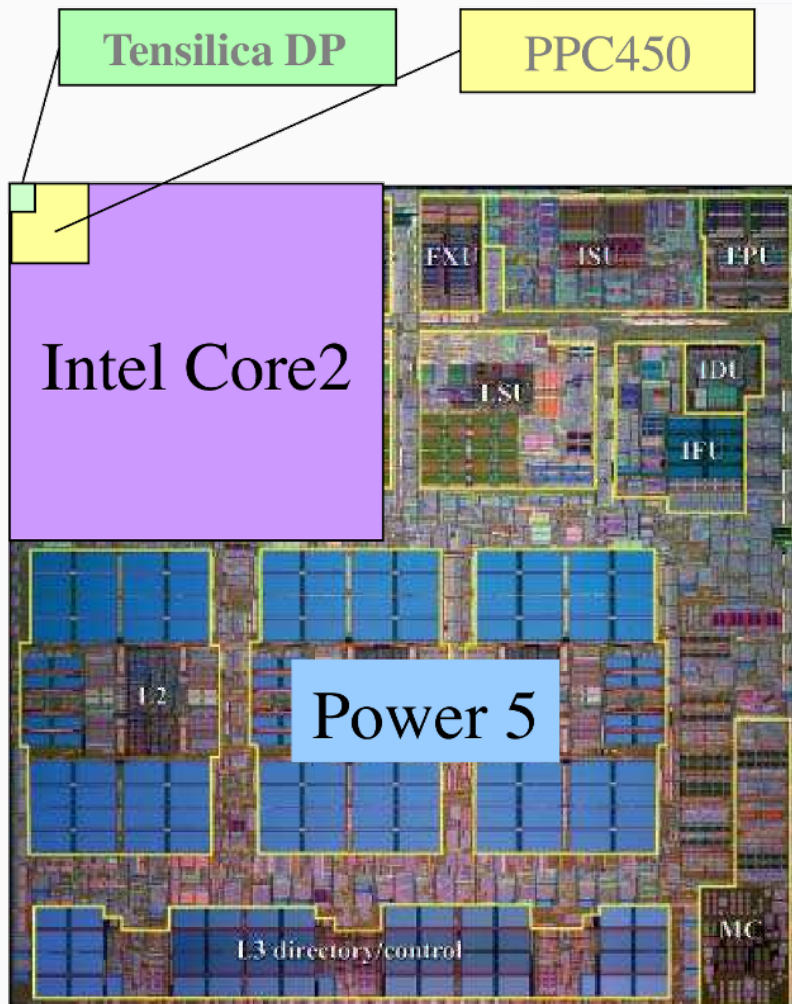


Image by John Shalf, LBNL

- IBM Power5: 120W @ 1900MHz
- Intel Core2 solo: 15W @ 1000MHz.
- IBM PPC 450 (Blue Gene): 0.625W @ 800MHz
- Tensilica XTensa (Moto Razor): 0.09W @ 600MHz

**400x** improvement in Flops per Watt!

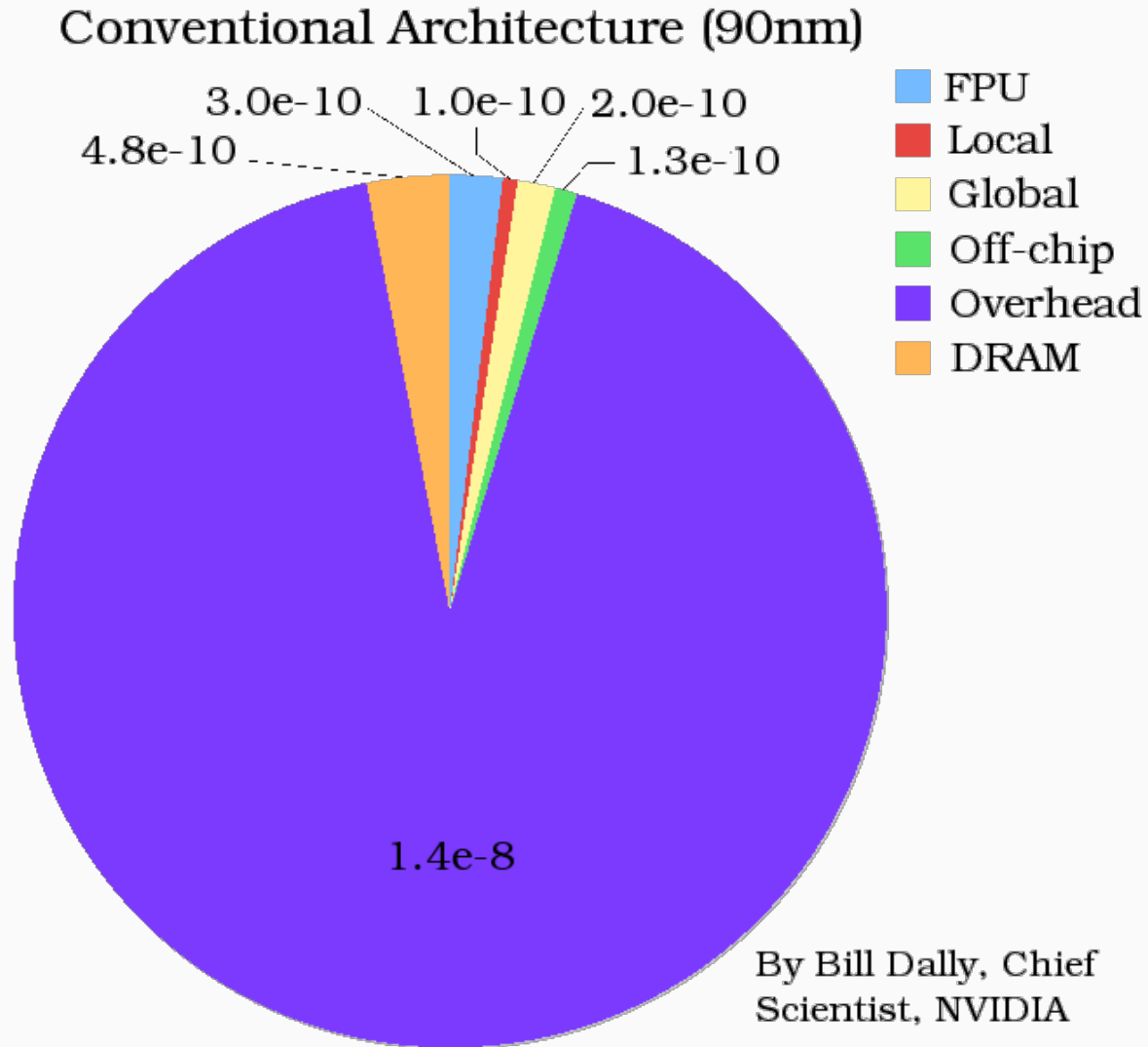
# *Rise of Many-core Systems*

## **Focus is on Flops per Watt:**

- Clock rates constant or decreasing.
- Use larger fraction of transistors for calculation, split into many “throughput” cores.
- Explicit memory hierarchy. Cache management now in software stack.  
RAM/node  $\uparrow$ , but RAM/core  $\downarrow$



# CPU Power Consumption



# *Rise of Many-core Systems*

## **Focus is on Flops per Watt:**

- Clock rates constant or decreasing.
- Use larger fraction of transistors for calculation, split into many “throughput” cores.
- Explicit memory hierarchy. Cache management now in software stack. RAM/node  $\uparrow$ , but RAM/core  $\downarrow$
- Keep some traditional “low latency” cores around for coordination.
- Filesystem I/O even more of a bottleneck...

# *“Throughput” Processors*

- NVIDIA Fermi: 480 cores @ 700MHz
- ATI Radeon 5970: 3200 cores @ 725MHz
- Intel Many Integrated Core (MIC): re-brand of failed Larrabee platform...
- Goal is to use something closer to 25% of transistors for Flops.
- Requires fine-grained parallelism, explicit memory movement.

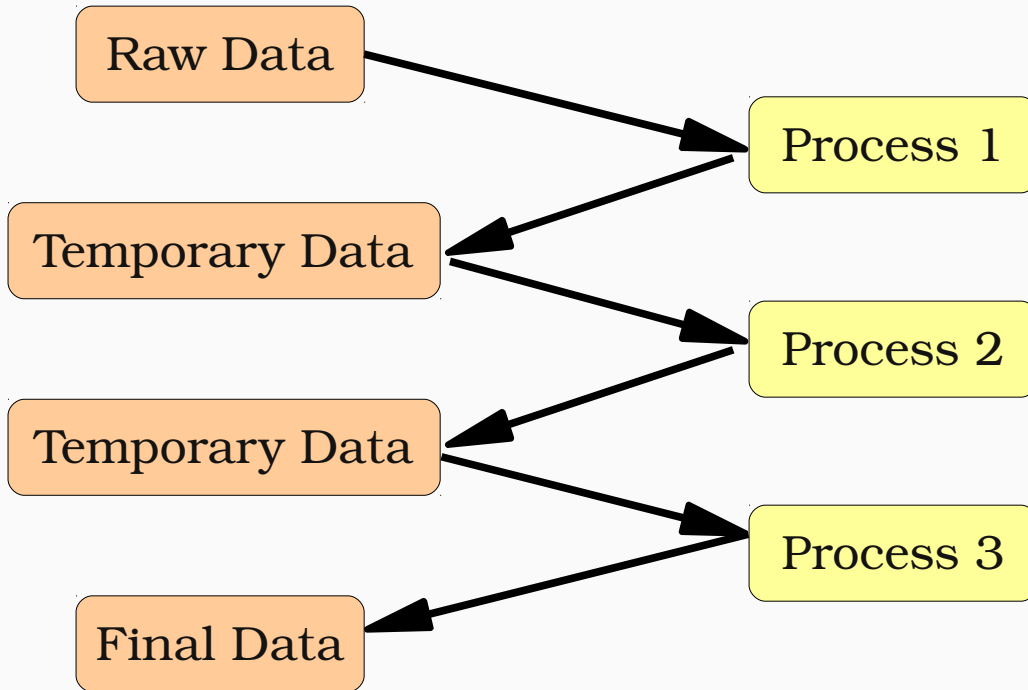
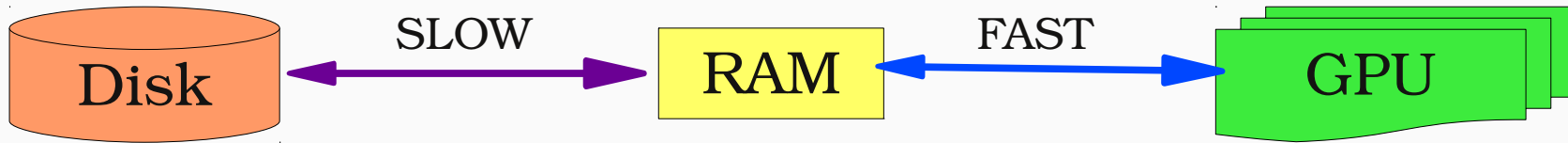
# *What does this mean for Astrophysics?*

- Astrophysical datasets are getting larger!
  - LSST: 15TB / day
  - Near-term CMB missions:  $O(100-1000)$  TB
- Systems in the very near future may have  $O(10)$  traditional cores and  $O(100-1000)$  throughput cores per node.
  1. Data movement can be more costly than calculations—minimize when possible.
  2. Determine what operations can be parallelized at the node level.
  3. Evaluate new tools as they become available.

# *Data Movement*

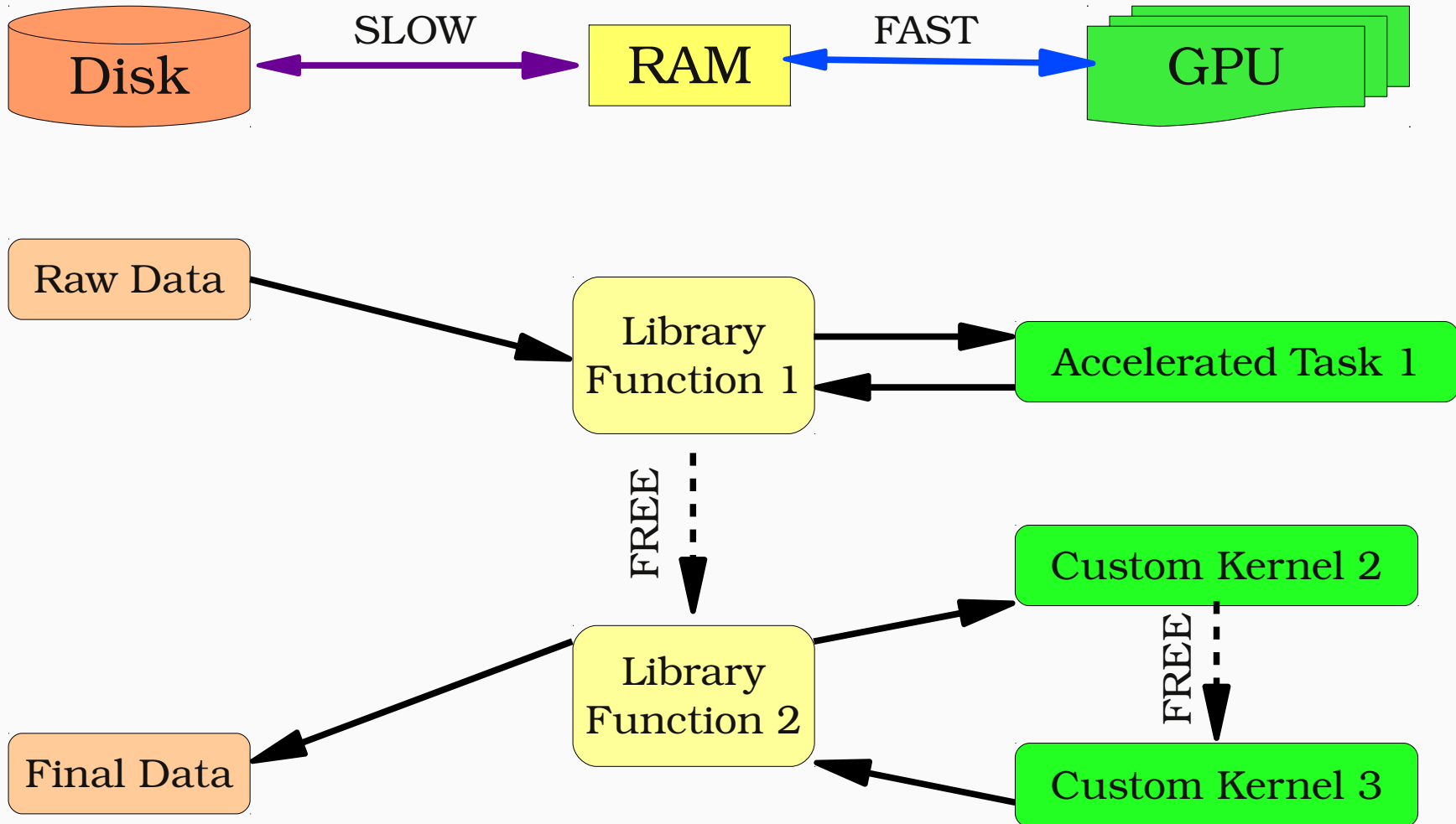
- Traditional paradigm:
  - Many small executables chained together
  - Write / read intermediate files
- This breaks down if:
  - I/O cost outpaces calculation AND
  - Overall runtime is unacceptably slow
- Movement to/from accelerators can also cancel benefit for some algorithms.

# Data Movement for “Chained Processes”



This is for playing Quake, right?

# Improved Data Movement



# *Parallelize Relevant Operations*

- Split processing based on independent data products (embarrassingly parallel work flows)
- 1D – time domain astrophysics:
  - vector math, FFTs, sparse matrix operations.
- 2D – image / map manipulation
  - Linear combinations, projections
  - convolution / filtering, spherical harmonic transforms
- 3D – data cube (spaxel/voxel) manipulations.



# *Parallelize Relevant Operations*

- Start by converting/switching low-level libraries
  - Likely to get some improvement without much work, e.g. FFT libraries.
- Only build custom code when needed- if data movement to/from card is dominant.
  - Use helper tools: PGI accelerator framework, MOAT (shameless plug!).

## *New Tools*

- We are faced with a huge diversity of platforms: GPUs/accelerators from different vendors, varying OS support.
- OpenCL: Unified interface to CPU/GPU devices, wide industry support.



# *Conclusions*

1. Start planning now for future hardware: will your code be ready for the cluster you purchase in 3 years?
2. Start testing new software tools that seem promising- what pieces of existing code are easy to parallize?
3. Will your future data volume overwhelm your current I/O patterns?